

# Higher spins of mixed symmetry

Andrea Campoleoni

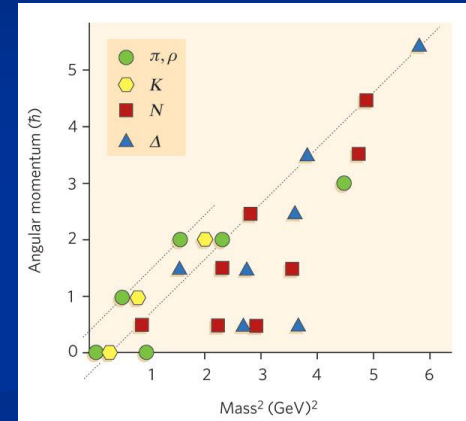
Albert Einstein Institute, Potsdam

A. C., D. Francia, J. Mourad and A. Sagnotti, 0810.4350, 0904.4447

For a review see A. C., 0905.1472, 0910.3155

# Why higher spins?

- Long standing open questions in field theory (see for instance Fierz-Pauli, 1939)
- Another setup where higher spins play a key role: **String Theory**
- String Theory was indeed designed to describe  $\infty$  higher-spin resonances
- What can we learn from String Theory?
- What can we add to the understanding of String Theory?
  - Example: String Theory was conjectured to be a broken phase of an higher-spin gauge theory. How to make precise this statement?



# String Theory and mixed symmetry

- String models are consistent in  $D > 4$
- If  $D > 4$  the little groups  $SO(D-1)$  or  $SO(D-2)$  admit more than a single Casimir operator
- Their irreps are characterized by sets of labels  $\{s_1, \dots, s_N\}$  with  $N \leq (D-1)/2$  or  $N \leq (D-2)/2$
- $N > 1 \rightarrow$  **Mixed symmetry** representations
- Mixed symmetry fields enter all String spectra
- Challenging setup for massless fields:
  - Symmetric fields ( $N=1$ ): Vasiliev full non-linear theory in AdS
  - No known extensions of the Vasiliev system for  $N > 1$

# Mixed-symmetry fields

- Free particles in  $D = 4$

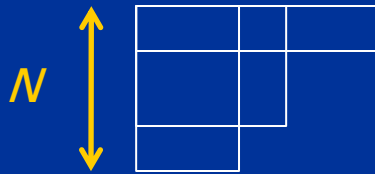
- Mass and spin

- Fully symmetric fields

$$\varphi_{\mu_1 \dots \mu_s} \quad \psi^\alpha_{\mu_1 \dots \mu_s}$$

- Free particles in  $D > 5$

- Irreps of  $SO(D-2)$



- Mixed-symmetry fields

$$\varphi_{\mu_1 \dots \mu_{s_1}, \nu_1 \dots \nu_{s_2}, \dots}$$

$$\psi^\alpha_{\underbrace{\mu_1 \dots \mu_{s_1}, \nu_1 \dots \nu_{s_2}, \dots}_N}$$

- In String Theory

$$\varphi_{\mu_1 \dots \mu_{s_1}, \nu_1 \dots \nu_{s_2}, \dots} \alpha_{-1}^{\mu_1} \dots \alpha_{-1}^{\mu_{s_1}} \alpha_{-2}^{\nu_1} \dots \alpha_{-2}^{\nu_{s_2}} \dots |0\rangle$$

- How to write field equations and Lagrangians?

# Outline

- Motivations
- **Constrained free theory for massless fields (metric-like)**
  - Non-Lagrangian field equations for Bose and Fermi fields
  - Lagrangians for Bose and Fermi fields
- **Comments on the role of the constraints**
- **Minimal unconstrained local theory for massless fields**
  - Non-Lagrangian field equations & auxiliary fields
  - Lagrangians & auxiliary fields
- Weyl-like symmetries
- Conclusions

# Massless fields & gauge symmetries

- $M = 0 \rightarrow$  gauge fields

- Gravity

$$\delta h_{\mu\nu} = \partial_{\mu} \xi_{\nu} + \partial_{\nu} \xi_{\mu}$$

- Fully symmetric fields

$$\delta \varphi_{\mu_1 \dots \mu_s} = \partial_{(\mu_1} \Lambda_{\mu_2 \dots \mu_s)}$$

- Mixed-symmetry fields

$$\delta \varphi_{\mu_1 \dots \mu_{s_1}, \nu_1 \dots \nu_{s_2}, \dots} = \partial_{(\mu_1} \Lambda_{\mu_2 \dots \mu_{s_1})}^{(1)}, \nu_1 \dots \nu_{s_2}, \dots + \partial_{(\nu_1} \Lambda_{\mu_1 \dots \mu_{s_1}, \nu_2 \dots \nu_{s_2})}^{(2)}, \dots + \dots,$$

- A good notation is needed...

- Hide the space-time indices

$$\varphi_{\mu_1 \dots \mu_{s_1}, \nu_1 \dots \nu_{s_2}, \dots} \rightarrow \varphi$$

- Select the “family” where the operators act

$$\partial_{(\mu_1} \Lambda_{\mu_2 \dots \mu_{s_1})}^{(1)}, \nu_1 \dots \nu_{s_2}, \dots \rightarrow \partial^1 \Lambda_1$$

- Gauge transformations:

$$\delta \varphi = \partial^i \Lambda_i$$

$$\delta \psi = \partial^i \epsilon_i$$

# Field equations

- An example: linearized gravity!

$$R_{\mu\nu} = \square h_{\mu\nu} - (\partial_\mu \partial \cdot h_\nu + \partial_\nu \partial \cdot h_\mu) + \partial_\mu \partial_\nu h^\lambda{}_\lambda = 0$$

- Labastida field equations

*Fronsdal, Fang (1978)*

*Labastida (1986)*

- Bose  $\mathcal{F} \equiv \square \varphi - \partial^i \partial_i \varphi + \frac{1}{2} \partial^i \partial^j T_{ij} \varphi = 0$
- Fermi  $\mathcal{S} \equiv i (\not{\partial} \psi - \partial^i \gamma_i \psi) = 0$

- Gauge variation

- Bose  $\delta \mathcal{F} = \frac{1}{2} \partial^i \partial^j \partial^k T_{(ij} \Lambda_{k)}$
- Fermi  $\delta \mathcal{S} = -\frac{i}{2} \partial^i \partial^j \gamma_{(i} \epsilon_{j)}$

- Gauge invariant field equations  $\rightarrow$  **constraints** on  $\Lambda_i, \epsilon_i$

# Again the gravity example...

- Non-Lagrangian field equation

$$R_{\mu\nu} = \square h_{\mu\nu} - (\partial_\mu \partial \cdot h_\nu + \partial_\nu \partial \cdot h_\mu) + \partial_\mu \partial_\nu h^\lambda{}_\lambda = 0$$

- “Natural ansatz” for the Lagrangian

$$\mathcal{L} = \frac{1}{2} h^{\mu\nu} (R_{\mu\nu} + k \eta_{\mu\nu} R_\lambda{}^\lambda),$$

- Gauge variation

$$\delta h_{\mu\nu} = \partial_\mu \xi_\nu + \partial_\nu \xi_\mu \quad \Rightarrow \quad \delta \mathcal{L} = - \xi^\mu (\partial^\lambda R_{\mu\lambda} + k \partial_\mu R_\lambda{}^\lambda),$$

- Bianchi identity

$$\partial^\lambda R_{\mu\lambda} - \frac{1}{2} \partial_\mu R_\lambda{}^\lambda = 0 \quad \Rightarrow \quad k = -\frac{1}{2}$$



# Lagrangians for arbitrary fields

## ■ Bianchi identities

■ Bose 
$$\partial_i \mathcal{F} - \frac{1}{2} \partial^j T_{ij} \mathcal{F} = -\frac{1}{12} \partial^j \partial^k \partial^l T_{(ij} T_{kl)} \varphi$$

■ Fermi 
$$\partial_i \mathcal{S} - \frac{1}{2} \partial^j \mathcal{S}_i - \frac{1}{2} \partial^j T_{ij} \mathcal{S} - \frac{1}{6} \partial^j \gamma_{ij} \mathcal{S} = \frac{i}{6} \partial^j \partial^k T_{(ij} \gamma_{k)} \psi$$

## ■ They lead to the constraints

$$T_{(ij} T_{kl)} \varphi = 0 \qquad T_{(ij} \gamma_{k)} \psi = 0$$

## ■ Natural ansatz for the Lagrangians

$$\mathcal{L}_{\text{Bose}} = \frac{1}{2} \langle \varphi, \sum_{p=0}^N k_p \eta^p \mathcal{F}^{[p]} \rangle$$

*Labastida (1986)*

$$\mathcal{L}_{\text{Fermi}} = \frac{1}{2} \langle \bar{\psi}, \sum_{p,q=0}^N k_{p,q} \eta^p \gamma^q (\gamma^{[q]} \mathcal{S}^{[p]}) \rangle + \text{h.c.}$$

*A.C., Francia, Mourad, Sagnotti (2009)*

# Constrained Lagrangians

## ■ Bose fields

*Labastida (1989)*

$$\mathcal{L}_{\text{Bose}} = \frac{1}{2} \langle \varphi, \sum_{p=0}^N \frac{(-1)^p}{p!(p+1)!} \eta^p \mathcal{F}^{[p]} \rangle$$

## ■ Fermi fields

*A.C., Francia, Mourad, Sagnotti (2009)*

$$\mathcal{L}_{\text{Fermi}} = \frac{1}{2} \langle \bar{\psi}, \sum_{p,q=0}^N \frac{(-1)^{p+\frac{q(q+1)}{2}}}{p!q!(p+q+1)!} \eta^p \gamma^q (\gamma^{[q]} \mathcal{S}^{[p]}) \rangle + \text{h.c.}$$

## ■ Constrained because...

$$\delta \mathcal{F} = \frac{1}{2} \partial^i \partial^j \partial^k T_{(ij} \Lambda_{k)}$$

$$\delta \mathcal{S} = -\frac{i}{2} \partial^i \partial^j \gamma_{(i} \epsilon_{j)}$$

$$\partial_i \mathcal{F} - \frac{1}{2} \partial^j T_{ij} \mathcal{F} = -\frac{1}{12} \partial^j \partial^k \partial^l T_{(ij} T_{kl)} \varphi$$

$$\partial_i \mathcal{S} - \frac{1}{2} \not{\partial} \mathcal{S}_i - \frac{1}{2} \partial^j T_{ij} \mathcal{S} - \frac{1}{6} \partial^j \gamma_{ij} \mathcal{S} = \frac{i}{6} \partial^j \partial^k T_{(ij} \gamma_{k)} \psi$$

# Comments on the constraints

- These constraints should not obstruct a possible non-linear deformation of the free theory

- The Vasiliev theory contains constrained fields

*Fradkin, Vasiliev (1987)*

*Vasiliev (1990-2010)*

- Still, they are quite unsatisfactory...

- No constraints in the massless limit of free String Field Theory

*Henneaux, Teitelboim (1988), Sagnotti, Tsulaia (2004)*

- The constrained theory cannot be related to the linearized curvatures of de Wit and Freedman

*De Wit, Freedman(1980)*

- Higher spin geometry:

$$\delta \Gamma_{\lambda, \mu\nu} = \partial_{\mu} \partial_{\nu} \xi_{\lambda}$$

$$\delta \Gamma^{(k)}_{\lambda_1 \dots \lambda_k, \mu_1 \dots \mu_s} = \partial_{(\mu_1} \dots \partial_{\mu_{k+1}} \Lambda_{\mu_{k+2} \dots \mu_s)} \lambda_1 \dots \lambda_k$$

- It is possible to use the HS curvatures to obtain **non-local** Lagrangians describing the correct **local** dynamics

*Francia, Sagnotti, Mourad (2002-2010), Bekaert, Boulanger (2002), De Medeiros, Hull (2003)*

# Unconstrained theory

- Key steps in constructing the constrained theory
  - Gauge invariance → kinetic tensors & constraints on the gauge parameters
  - Bianchi identities → constraints on the fields
  - Traces of the Bianchi identities → uniquely fix the Lagrangians
- In the unconstrained theory
  - Gauge invariance → auxiliary compensator fields
  - Bianchi identities → Lagrange multipliers
  - Traces of the Bianchi identities → fix the Lagrangians up to field redefinitions

# Gauge invariance & auxiliary fields

## ■ Gauge variation of the kinetic tensors

- Bose  $\delta \mathcal{F} = \frac{1}{2} \partial^i \partial^j \partial^k T_{(ij} \Lambda_{k)}$  *Francia, Sagnotti (2002, 2005)*  
*Francia, Mourad, Sagnotti (2007)*
- Fermi  $\delta \mathcal{S} = -\frac{i}{2} \partial^i \partial^j \gamma_{(i} \epsilon_{j)}$  *A.C., Francia, Mourad, Sagnotti (2008)*

## ■ Compensator fields

- Bose  $\delta \Phi_i = \Lambda_i \rightarrow \mathcal{A} = \mathcal{F} - \frac{1}{6} \partial^i \partial^j \partial^k T_{(ij} \Phi_{k)}$
- Fermi  $\delta \Psi_i = \epsilon_i \rightarrow \mathcal{W} = \mathcal{S} + \frac{i}{2} \partial^i \partial^j \gamma_{(i} \Psi_{j)}$

- Eliminating the auxiliary fields one recovers the non-local field equations of Francia and Sagnotti *Francia (2010)*

# Bianchi identities & auxiliary fields

## ■ Unconstrained Bianchi identities

$$\partial_i \mathcal{A} - \frac{1}{2} \partial^j T_{ij} \mathcal{A} = -\frac{1}{12} \partial^j \partial^k \partial^l T_{(ij} T_{kl)} (\varphi - \partial^m \Phi_m) \quad \mathcal{C}_{ijkl}$$

$$\partial_i \mathcal{W} - \frac{1}{2} \partial^j \mathcal{W}_i - \frac{1}{2} \partial^j T_{ij} \mathcal{W} - \frac{1}{6} \partial^j \gamma_{ij} \mathcal{W} = \frac{i}{6} \partial^j \partial^k T_{(ij} \gamma_{k)} (\psi - \partial^m \Psi_m) \quad \mathcal{Z}_{ijk}$$

## ■ Lagrange multipliers to compensate the remainder

$$\mathcal{L}_{\text{Bose}} = \frac{1}{2} \langle \varphi - \partial^m \Phi_m, \sum_{p=0}^N \frac{(-1)^p}{p!(p+1)!} \eta^p \mathcal{A}^{[p]} \rangle + \frac{1}{8} \langle \beta_{ijkl}, \mathcal{C}_{ijkl} \rangle$$

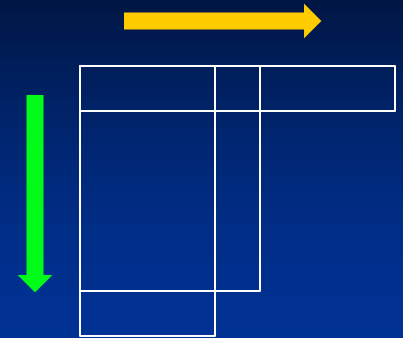
$$\mathcal{L}_{\text{Fermi}} = \frac{1}{2} \langle \bar{\psi} - \partial^m \bar{\Psi}_m, \sum_{p,q=0}^N \frac{(-1)^{p+\frac{q(q+1)}{2}}}{p!q!(p+q+1)!} \eta^p \gamma^q (\gamma^{[q]} \mathcal{W}^{[p]}) \rangle$$

$$+ \frac{1}{4} \langle \bar{\lambda}_{ijk}, \mathcal{Z}_{ijk} \rangle + \text{h.c.}$$

# Structure of the Lagrangians

- Two peculiar features:

- Compensators and Lagrange multipliers
- Higher traces or  $\gamma$ -traces of the kinetic tensor in the Lagrangians



$$\mathcal{L}_{\text{Bose}} = \frac{1}{2} \langle \varphi - \partial^m \Phi_m, \sum_{p=0}^N \frac{(-1)^p}{p!(p+1)!} \eta^p \mathcal{A}^{[p]} \rangle + \frac{1}{8} \langle \beta_{ijkl}, \mathcal{C}_{ijkl} \rangle$$

$$\mathcal{L}_{\text{Fermi}} = \frac{1}{2} \langle \bar{\psi} - \partial^m \bar{\Psi}_m, \sum_{p,q=0}^N \frac{(-1)^{p+\frac{q(q+1)}{2}}}{p!q!(p+q+1)!} \eta^p \gamma^q (\gamma^{[q]} \mathcal{W}^{[p]}) \rangle + \frac{1}{4} \langle \bar{\lambda}_{ijk}, \mathcal{Z}_{ijk} \rangle + \text{h.c.}$$

Higher spins

Mixed symmetry

# Equations of motion

- Field equations for bosons

$$\sum_{p=0}^N \frac{(-1)^p}{p!(p+1)!} \eta^p \mathcal{F}^{[p]} + \frac{1}{2} \eta^{ij} \eta^{kl} \mathcal{B}_{ijkl} = 0$$

$$T_{(ij} T_{kl)} \mathcal{F} = 0$$

- Field equations for fermions

$$\sum_{p,q=0}^N \frac{(-1)^{p+\frac{q(q+1)}{2}}}{p!q!(p+q+1)!} \eta^p \gamma^q (\gamma^{[q]} \mathcal{S}^{[p]}) - \frac{1}{2} \eta^{ij} \gamma^k \mathcal{Y}_{ijk} = 0$$

$$T_{(ij} \gamma_k) \mathcal{S} = 0$$

- Are they equivalent to  $\mathcal{F} = 0$  and  $\mathcal{S} = 0$  ?

**YES if  $D \geq 2(N+1)$  but...**



# Possible pathologies

- Recall 2D gravity

$$R_{\mu\nu} - \frac{1}{2} \eta_{\mu\nu} R = 0 \quad \not\Rightarrow \quad R_{\mu\nu} = 0$$

- In general what can happen?

- $\mathcal{F}$  or  $\mathcal{S}$  do not vanish but the Lagrangians do

- Two-dimensional graviton & gravitino
    - Two-column bosons and single-column fermions in low enough space-time dimensions

- The Lagrangians possess more symmetries with respect to the non-Lagrangian field equations

- Some classes of mixed-symmetry bosons and fermions for  $D < 2(N+1)$

- Less equations but new symmetries

$$\delta \mathcal{F} = \eta^{ij} \Theta_{ij}$$

$$\delta \mathcal{S} = \gamma^i \Theta_i$$

# The classification

A.C., Francia, Mourad, Sagnotti (2008-2009)

- $[T_{ij}, \eta^{kl}] = \frac{D}{2} \delta_i^{(k} \delta_j^{l)} + \frac{1}{2} \left( \delta_i^{(k} S^l)_{j} + \delta_j^{(k} S^l)_{i} \right)$ .
- $S^i_j$  operators:  $S^1_2 \varphi \rightarrow \varphi(\mu_1 \dots \mu_{s_1}, \mu_{s_1+1}) \nu_1 \dots \nu_{s_2-1}, \dots$

- Considering  $\delta \mathcal{F} = \eta^{ij} \Theta_{ij}$  and  $\delta \mathcal{S} = \gamma^i \Theta_i$  leads to

- Bose  $(D-2) \Theta_{ij} + S^k_{(i} \Theta_{j)k} = 0$

- Fermi  $(D-2) \Theta_i + 2 S^j_i \Theta_j = 0$

- Classification of Weyl-like symmetries



eigenvalue problems for the  $S^i_j$  operators

$$[S^i_j, S^k_l] = \delta_j^k S^i_l - \delta_l^i S^k_j$$

← gl(N) algebra

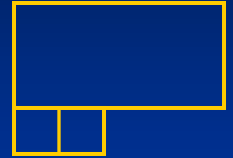
# Weyl invariant models

*A.C., Francia, Mourad, Sagnotti (2009)*

## ■ Examples of solutions of the previous conditions

■ Bose  $\{2^{N-1}, 2\}$  – projected fields in  $D = 2N$

■ Fermi  $\{2^{N-1}, 1\}$  – projected fields in  $D = 2N$



## ■ The Weyl-like shifts also preserve

■ Bose  $T_{(ij} T_{kl)} \mathcal{F} = 0$

■ Fermi  $T_{(ij} \gamma_{k)} \mathcal{S} = 0$

## ■ Other Weyl-invariant models

■ Bose  $\{2^p, 1^q\}$  – projected fields in  $D \leq 2p+q$

■ Fermi  $\{1^q\}$  – projected fields in  $D \leq 2q$



## ■ In all these cases **no local d.o.f. are involved**

# Summary & outlook

- Simple guidelines
  - Gauge invariance
  - Bianchi identities
- New results
  - Completion of the constrained theory for **Fermi** fields
    - Check of the propagating d.o.f.
    - Lagrangians
  - Unconstrained Lagrangians for **Bose** and **Fermi** fields
  - Weyl-like symmetries
- Perspectives
  - Extension to (A)dS, Supersymmetry
  - Links with String Theory
  - Interactions...